# Advanced Statistical Physics

Lecture 2 : Fundamental Concepts - Part II

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- Please, write down your expectations from the course.
- $\cdot$  Typos of the course
- Recommendations are always welcome.

- Motivation for studying of statistical physics
- Review on thermodynamics
- A summary on thermodynamics through statistical relations

# A crash course on statistical mechanics.

- 1. The heroes
- 2. Our lord and savior "Z"
- 3. Statistics of Ideal Gases

"With thermodynamics, one can calculate almost everything crudely; with kinetic theory, one can calculate fewer things, but more accurately; and with statistical mechanics, one can calculate almost nothing exactly." Eugene Paul Wigner The heroes

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An ensemble that follows the probability distribution (6) to describe isolated systems, is called *the microcanonical ensemble*.

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Question : What is the probability  $P_r$  of finding system A in any one particular microstate r of energy  $E_r$ ?

Answer:

$$P_r = \frac{e^{-\beta E_r}}{\sum_r e^{-\beta E_r}}$$

This expression is central in statistical physics, the exponential factor  $e^{-\beta E_r}$  is called the **"Boltzmann factor"** and the probability distribution is the **"canonical distribution"**. Any ensemble describing systems interacting with a heat bath characterized with a temperature *T*, following the canonical distribution, is called the **"canonical ensemble"**.

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What is the mean magnetic momentum  $\bar{\mu}_{H}$  of such an atom ?

#### Application : Paramagnetism

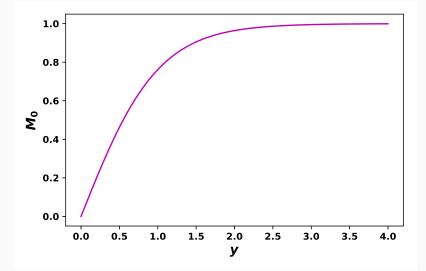


Figure 1: Magnetization with respect to y.

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A particular example is **the grand canonical ensemble**, in which X = N, meaning that the system can exchange the energy *E* and the particles *N*. The parameter  $\beta$  is the temperature of the ensemble, and  $\alpha = \frac{-\mu}{kT}$ , where  $\mu$  is **the chemical potential** of the reservoir.

# Our lord and savior "Z"

Statistical physics has a dear friend that we've been trying our best to hide so far. This friend, is our greatest ally in order to grasp the physical situations of our interest. Statistical physics has a dear friend that we've been trying our best to hide so far. This friend, is our greatest ally in order to grasp the physical situations of our interest.

$$Z = \sum_{r} e^{-\beta E_{r}}.$$
 (1)

*"Z"* is a sum over states, called **the partition function**. *Z* is used as symbol for the parition function due to its name in German *"Zustandssumme"* 

### Application : The one dimensional harmonic oscillator

The 1D harmonic oscillator is a famous model in physics, it is used to describe any physical situation where the potential energy is a quadratic function of the coordinate, such as the vibration of molecule around its position of equilibrium.

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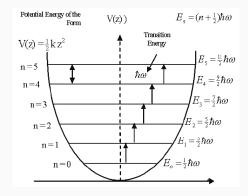


Figure 2: Energy levels of the 1D harmonic oscillator

### Application : The one dimensional harmonic oscillator

$$Z = \frac{1}{2\sinh(-\beta\frac{\hbar\omega}{2})}$$
$$\bar{E} = -\frac{\partial\ln Z}{\partial\beta} = -\frac{\hbar\omega}{2}\coth(-\beta\frac{\hbar\omega}{2}),$$
$$S = k\left(\ln Z + \beta\bar{E}\right) = k\left[-\ln\left(2\sinh\left(\frac{\hbar\omega}{2kT}\right)\right) + \frac{\hbar\omega}{2kT}\coth\left(\frac{\hbar\omega}{2kT}\right)\right],$$
$$F = -kT\ln Z = kT\ln\left(2\sinh\left(\frac{\hbar\omega}{2kT}\right)\right)$$

## Statistics of Ideal Gases

Consider a gas of identical particles characterized by a volume V, temperature T and a number of possible quantum states R. Each particle can be in quantum state labeled by r with an energy  $\epsilon_r$  and the number of particles occupying the state r is  $n_r$ .

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## Statistical distribution functions

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The mean number of particles in a state s is :

$$\bar{n}_{\rm s} = -\frac{1}{\beta} \frac{\partial \ln Z}{\partial \epsilon_{\rm s}}$$

The Maxwell-Boltzmann statistics takes into account the distinguishability of particles, as a result the partition function given by :

$$Z = \sum_{n_1, n_2, \dots} \frac{N!}{n_1! n_2! \dots} e^{-\beta(n_1 \epsilon_1 + n_2 \epsilon_2 + \dots)},$$

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The mean number of particles in a state s is :

$$\bar{n}_{\rm S} = N \frac{e^{-\beta\epsilon_{\rm S}}}{\sum_{\rm r} e^{-\beta\epsilon_{\rm r}}}.$$

This distribution is called : the Maxwell-Boltzmann distribution.

Now, we will consider a *special* case of the Bose-Einstein statistics where the total number *N* is not fixed. The partition function is given by :

$$Z=\sum_{R}e^{-\beta(n_1\epsilon_1+n_2\epsilon_2+\dots)},$$

then :

$$\ln Z = -\sum_{r} \ln \left(1 - e^{-\beta \epsilon_r}\right).$$

The mean number of particles in the state s :

$$\bar{n}_{\rm s}=\frac{1}{e^{\beta\epsilon_{\rm s}}-1}.$$

This is called the "Planck distribution".

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(2)  
$$= \frac{\sum_{n_{s}} n_{s} e^{-\beta n_{s}\epsilon_{s}} \sum_{n_{1}...n_{2}}^{(s)} e^{-\beta(n_{1}\epsilon_{1}+n_{2}\epsilon_{2}+...)}}{\sum_{n_{s}} e^{-\beta n_{s}\epsilon_{s}} \sum_{n_{1}...n_{2}}^{(s)} e^{-\beta(n_{1}\epsilon_{1}+n_{2}\epsilon_{2}+...)}}.$$
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(3)  
$$\bar{n}_{s} = \frac{1}{e^{\alpha+\beta\epsilon_{s}}+1}.$$
(4)

This is called the Fermi-Dirac distribution.

Following the same reasoning, but this time the sum ranges over all values of the numbers  $n_1, n_2, \ldots$  such that  $n_r = 0, 1, 2, 3 \ldots$ 

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$$\bar{n}_{\rm s} = \frac{1}{e^{(\beta\epsilon_{\rm s}+\alpha)} - 1}.$$
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 $\alpha = -\beta\mu$  is again the chemical potential and the formula (??) represents the "Bose-Einstein statistics".

## Remark : The ground state

In the limit  $T \rightarrow 0$ , BE and FD statistics behaves differently.

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• For BE statistics :

It is fine to have multiple particles in the same state, so to reach the lowest energy of the whole gas (at  $T \rightarrow 0$ ), all the particles need to be placed in their lowest-lying state of energy  $\epsilon_1$ . In the limit  $T \rightarrow 0$ , BE and FD statistics behaves differently.

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• For FD statistics :

One particle per state. Then, in order to reach the lowest energy state of the whole gas, one need to populate all the single-particle states starting from the state of the lowest energy  $\epsilon_1$  until all the particles are accommodated. The gas as a whole is in its state of lowest energy, but there are particles that have a very high energy compared to  $\epsilon_1$ .

As  $\beta \to 0$  in FD and BE statistics, large energies  $\epsilon_r$  contribute to the sum. To prevent this sum from exceeding *N*,  $\alpha$  must become large enough to that each  $\epsilon_r$  is sufficiently small. That is  $e^{\alpha + \beta \epsilon_r} \gg 1$ , then FD and BE statistics reduces to :

$$\bar{n}_r = N rac{e^{-eta \epsilon_r}}{\sum_r e^{-eta \epsilon_r}}.$$

Then, in the classical limit BE and FD statistics reduces to MB statistics.

- Chapter 6, 7 and 9 of "F. Reif Fundamentals of Statistical and Thermal Physics."
- Chapter 5, 6, 7 and 9 of "C. Ngo H. Ngo Physique statistique".

## Next lecture is dedicated to exercises.

## Thank you for your attendance and your attention.